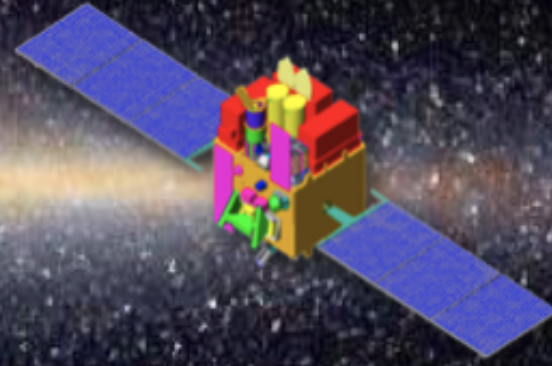


ASTROSAT

Cadmium Zinc Telluride Imager



CZTI Observations of GRBs

Vidushi Sharma

**5th July, 2017
TIFR-NCRA GRB workshop**

Overview

- *AstroSat: CZTI*
- *Detection of GRBs*
- *Lightcurves and spectrum*
- *Distribution on sky*
- *GRB statistics in 1 year*
- *Fluence distribution*
- *T_{90} distribution*

CZTI: Cadmium Zinc Telluride Imager

Hard X-ray detector

CZTI consists of four quadrants,
16 pixelated detectors each.

Imaging Method:

Coded Aperture Mask(CAM): It casts
a shadow on the detector when
illuminated by a source.

Energy range: **100 keV to < 250 keV**

Effective Area $\sim 500 \text{ cm}^2$

Thickness = 5 mm

CZTI Veto Detector(Cesium Iodide),
covers energy range **200keV to 500 keV.**

**CZT + Veto detectors together present
large field of view(FoV) comprising of the
whole sky excluding earth occultation.**



GRB detection by CZTI

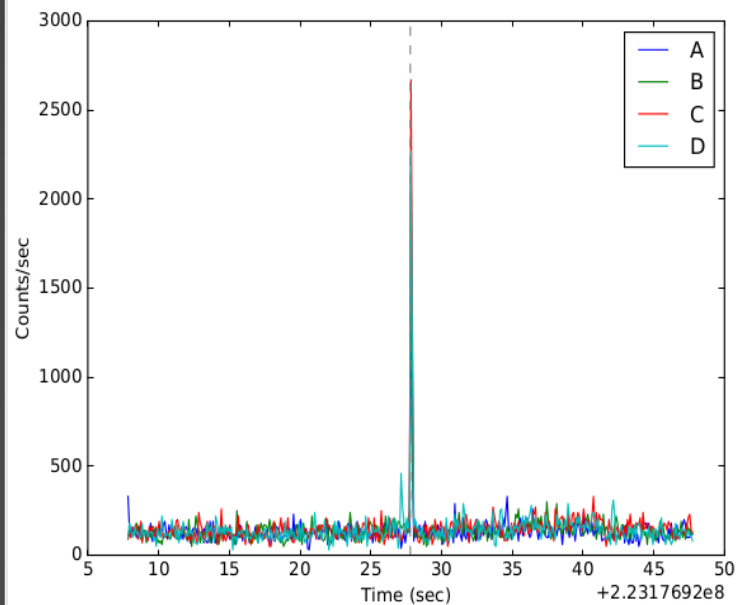
- All reported GRBs by GCN circulars through BAT or GBM or any other GRB detector are searched in the CZTI lightcurves with different binnings by the visual inspection.
- The GRB detection, localization, T_{90} and lightcurves by Astrosat are provided on the AstroSat CZTI GRB archive page.
- Total GRB triggers : 384 (6th Oct.,2015 - 30th June, 2017)
we expect, GRBs occurrence > 900 (In whole sky w.r.t. BAT(~2 sr))

<i>Fermi/GBM</i>	<i>~231</i>
<i>Swift/BAT</i>	<i>149</i>

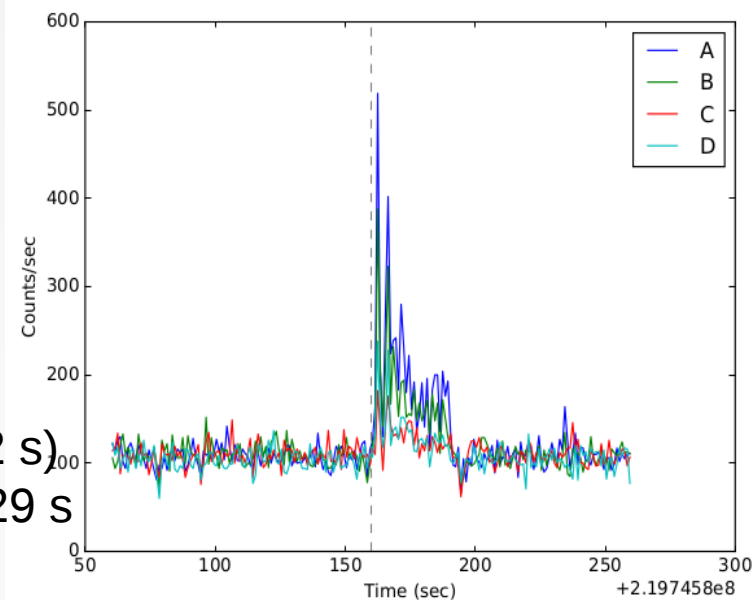
- Total GRBs detection by AstroSat:

<i>CZTI</i>	<i>121</i> <i>(sGRBs=11)</i>
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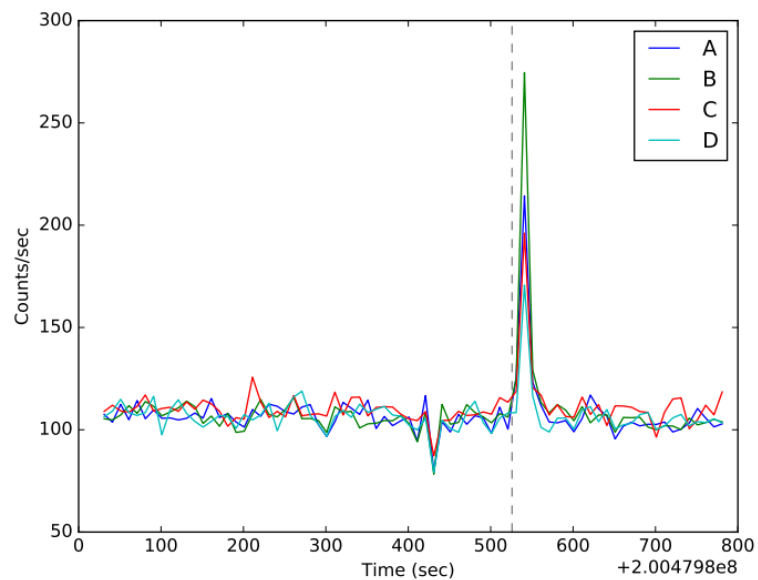
GRB lightcurves



Short GRBs (< 2 s)
GRB 170127C, 0.91 s

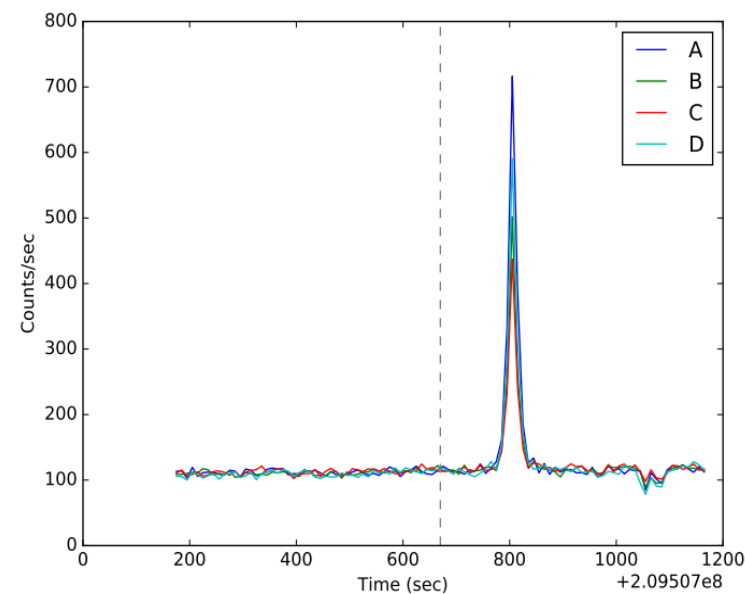


Long GRBs (> 2 s)
GRB 161218B, 29 s



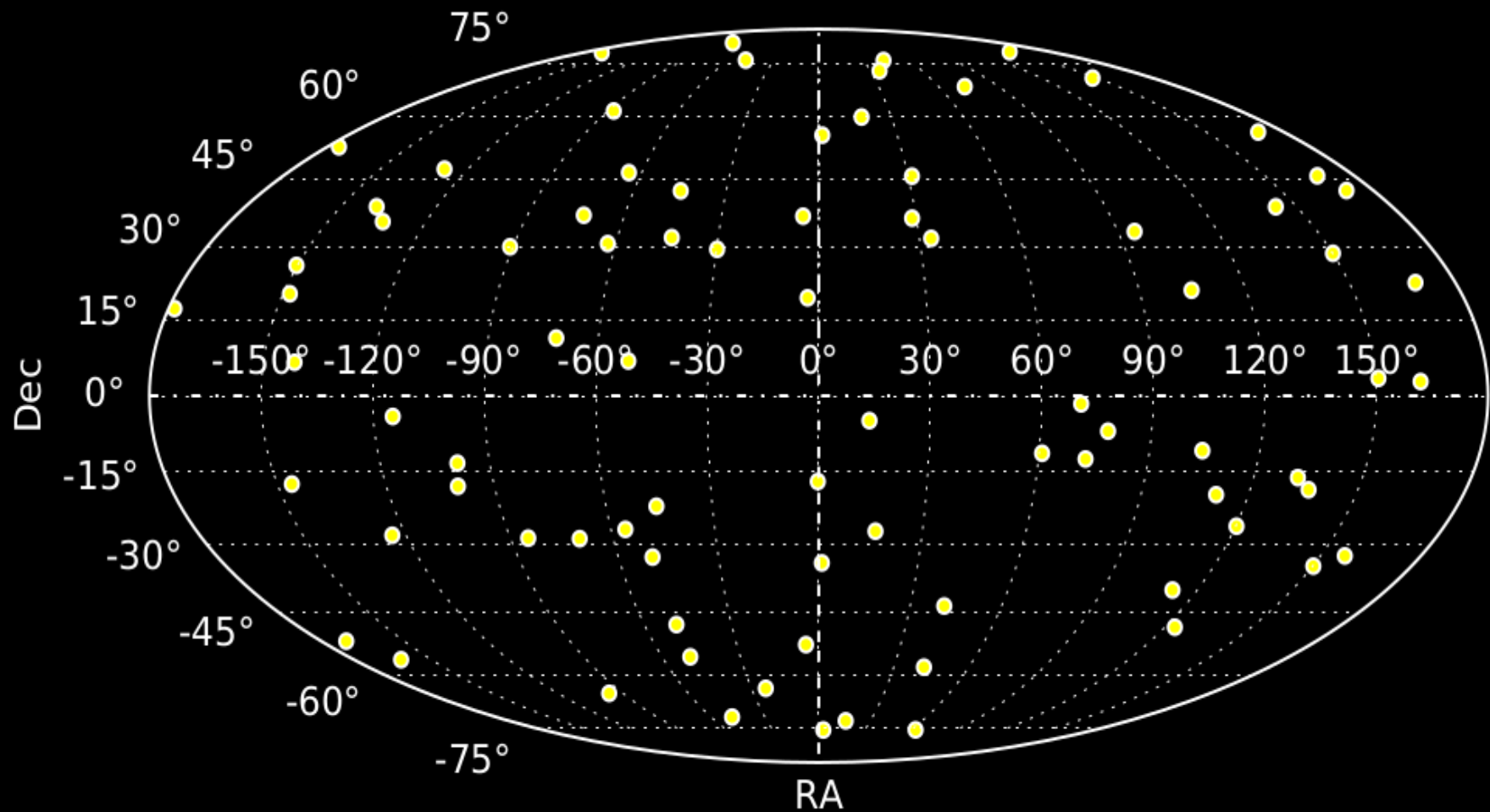
GRB 160509A

GRB 160821A



GRBs distribution on the sky: Isotropic

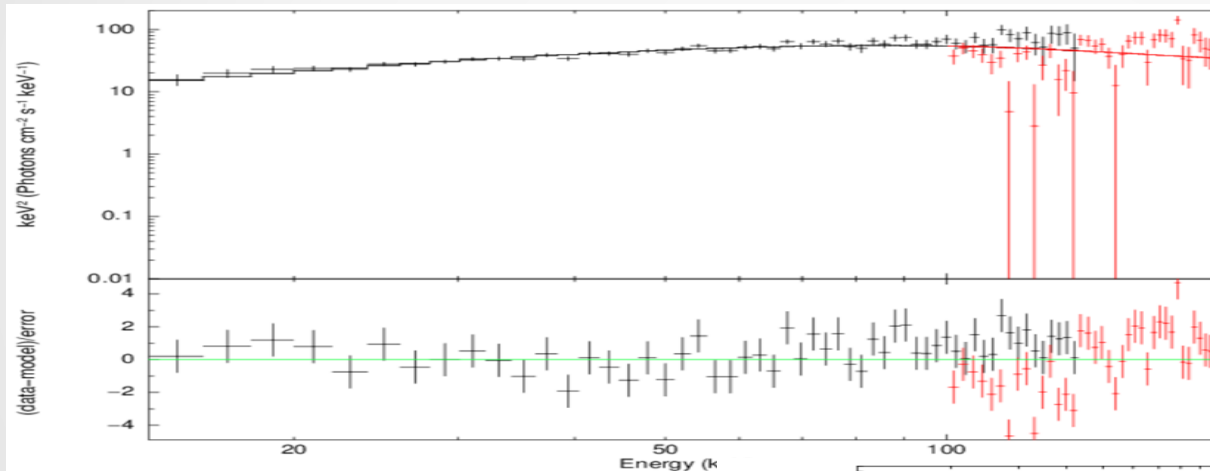
GRBs with known RA Dec = 82, till 31st Jan., 17



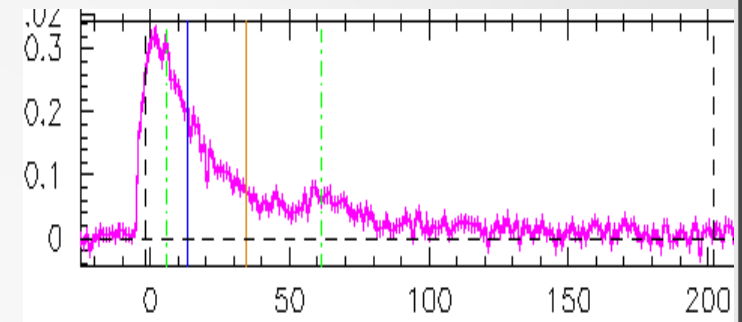
GRB 151006A Spectrum

First reported by Fermi/GBM, followed by Swift/BAT and Astrosat/CZTI.
 $T_{90}=84$ s

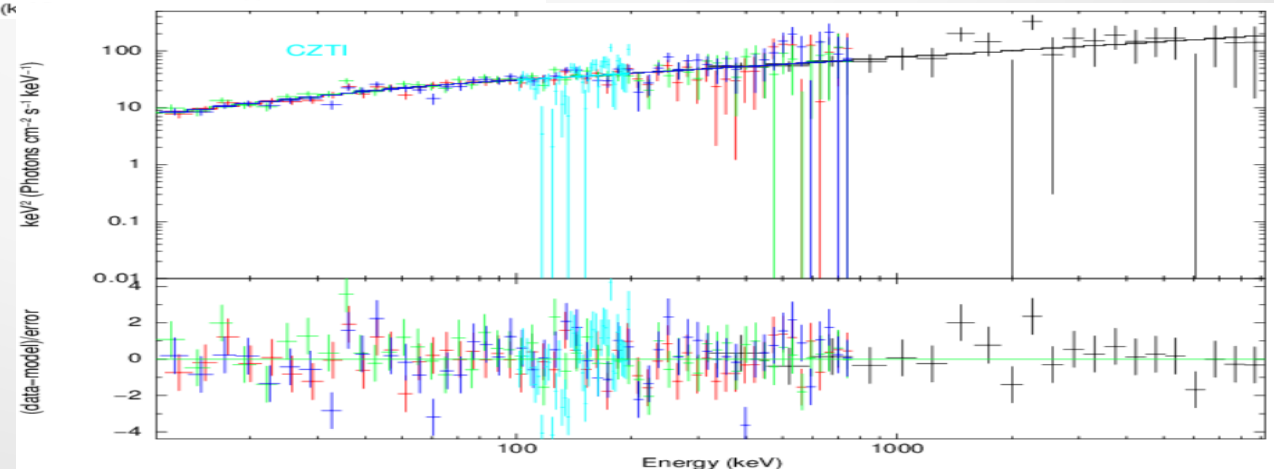
$\Theta_x, \Theta_y = 34.35, 58.87$



**GBM+CZT fitted with
Band function model**



**BAT+CZT fitted with
Band Function Model**



GRB statistics in 1 year

6th Oct 2015 - 6th Oct 2016

Total detected GRBs = 249

we expect > 530 GRBs occurrence in whole(4π sr) sky w.r.t. BAT(~ 2 sr)

Fermi/GBM	158
Swift/BAT	88
AstroSat/CZTI CZT, Veto	76 48, 73

Overlap in detections

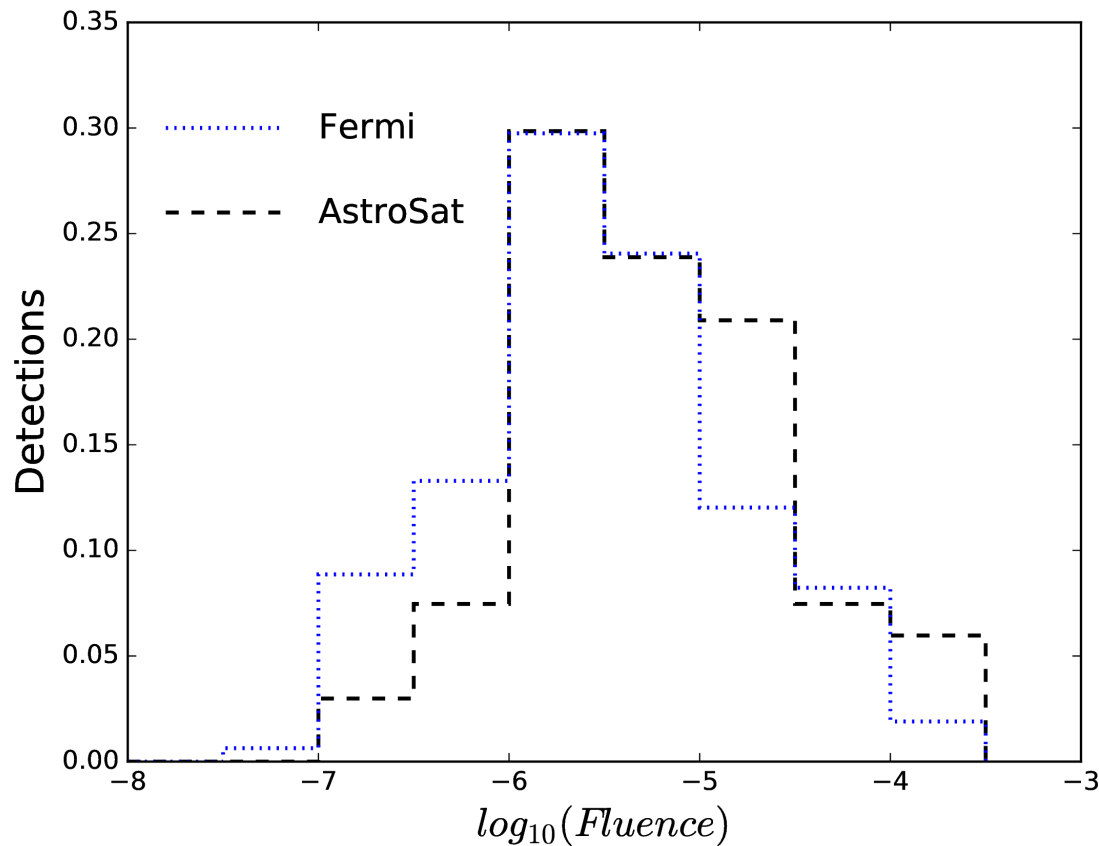
CZTI+BAT	Expected= $88 \times 55\% \sim 48$ Observed=25
CZTI+GBM	Expected= $158 \times 55\% \sim 87$ Observed=54
AstroSat SAA+Data Gap Earth Occultation Total	$\sim 20.5\%$ $\sim 25\%$ (almost comparable to fermi) $\sim 45\%$

GRBs not overlapping With FERMI & SWIFT:

GRB 160525C: CALET, GRB 160721A: CALET, GRB 160804B: CALET

GRB 160915A: Lomonosov, GRB 160917B: Lomonosov

Fluence Distribution

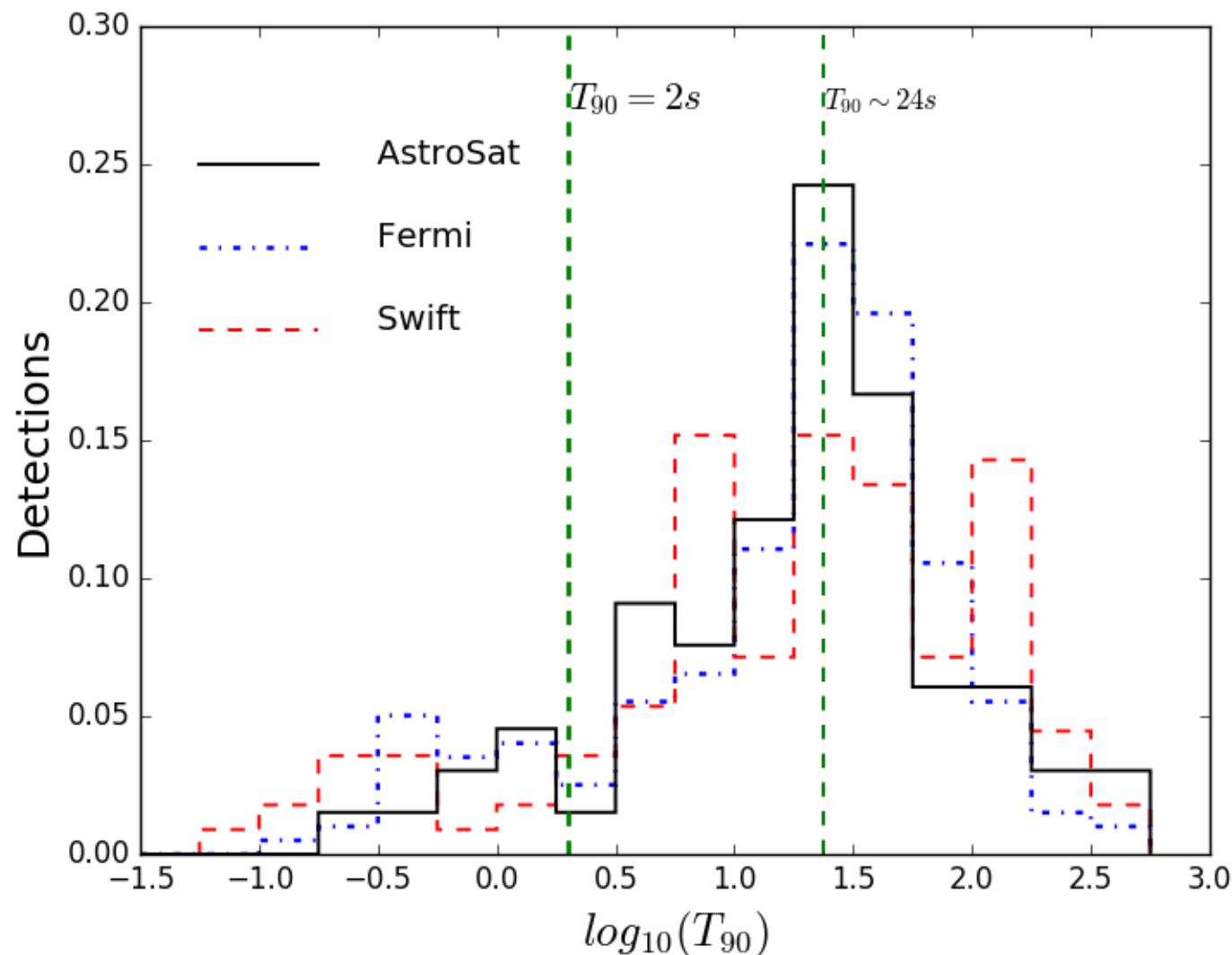


Fluence distributions for Fermi GBM and Astrosat CZTI(in Fermi units).
The distributions are normalized to unity.

**It detects relatively lesser number of bursts with
fluence less than 10^{-6} ergs cm^{-2} .**

Durations

T_{90} distribution for Fermi GBM, Swift BAT & AstroSat CZTI



Thus CZTI is sensitive to GRBs of all durations.

The bimodality distribution will be more clearer with further detections.

Under Progress

- More GRBs detections with automated GRB pipeline
- For CZTI spectrum and parameters: Mass modelling

THANK YOU!!!

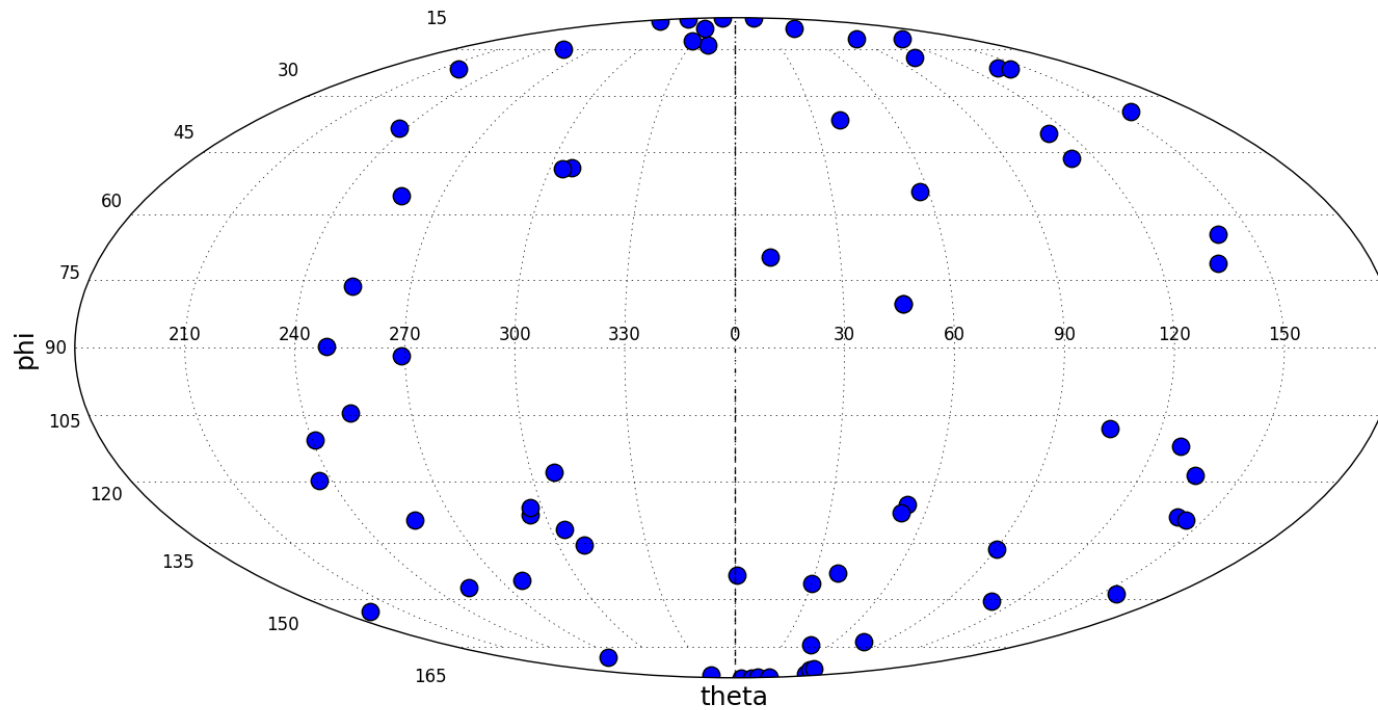
1)

CZTI	121
CZT	83
Veto	112

GRBs distribution on the sky: theta, phi

Using conventional division of $T_{90}=2s$ b/w short & long duration bursts, we find 7(9%) out of 76 GRBs lie in short GRB.
Fermi(~15.8%) and swift(~15.9%)

theta, phi distribution: In 1 year



RA Dec distribution till 30th June, 2017

GRBs with known RA, Dec=103

